BRITISH GEOLOGICAL SURVEY

Hartland Observatory

Monthly Magnetic Bulletin

October 2005

05/10/HA

HARTLAND POINT









1. HARTLAND OBSERVATORY MAGNETIC DATA

1.1 Introduction

This bulletin is published to meet the needs of both commercial and academic users of geomagnetic data. Magnetic observatory data is presented as a series of plots of one-minute, hourly and daily values, followed by tabulations of monthly values, geomagnetic activity indices and reports of rapid variations. The operation of the observatory and presentation of data are described in the rest of this section.

Enquiries about the data should be addressed to:

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Tel: +44 (0) 131 667 1000 Fax: +44 (0) 131 668 4368 E-mail: o.baillie@bgs.ac.uk Internet: www.geomag.bgs.ac.uk

1.2 Position

Hartland Observatory, one of the three geomagnetic observatories operated and maintained in the UK by BGS, is situated on the NW boundary of the village of Hartland in North Devon. The observatory coordinates are:

Geographic: 50°59.7′N 355°31.0′E Geomagnetic: 53°53.7′N 080°09.4′E Height above mean sea level: 95 m

The geomagnetic co-ordinates are calculated using the 10th generation International Geomagnetic Reference Field at epoch 2005.5.

1.3 The Observatory Operation

1.3.1 GDAS

The observatory operates under the control of the Geomagnetic Data Acquisition System (GDAS), which was developed by BGS staff, installed in 2002, and became fully operational in January 2003. The system operates under the control of data acquisition software running on QNX computers, which control the data logging and communications.

There are two sets of sensors used for making magnetic measurements. A triaxial linear-core fluxgate magnetometer, manufactured by the Danish Meteorological Institute, is used to measure the variations in the horizontal (*H*) and vertical (*Z*) components of the field. The third sensor is oriented perpendicular to these, and measures variations, which are proportional to the changes in declination (*D*). Measurements are made at a rate of 1 Hz.

In addition to the fluxgate sensors there is a proton precession magnetometer making measurements of the absolute total field intensity (F) at a rate of 0.1Hz.

The raw unfiltered data are retrieved automatically via Internet connections to the BGS office in Edinburgh in near real-time. The fluxgate data are filtered to produce one-minute values using a 61-point cosine filter whilst the total field intensity samples are filtered using a 7-point cosine filter. These one-minute values are used to update the Geomagnetism Information and Forecast Service (GIFS), an on-line information system accessed via the World Wide Web at the address given in Section 1.1. GIFS also provides information on geomagnetic and solar activity.

1.3.2 Back-up Systems

There are two other fully independent identical systems, GDAS 2 and GDAS 3, operating at the observatory. The data from these are also processed in near real-time and used for quality control purposes. They can also be used to fill any gaps or replace any corrupt values in the primary system, GDAS 1.

1.4 Data Presentation

The data presented in the bulletin are in the form of plots and tabulations described in the following sections.

1.4.1 Summary magnetograms

Small-scale magnetograms are plotted which allow the month's data to be viewed at a glance. They are plotted 16 days a page and show the variations in D, H and Z. The scales are shown on the right-hand side of the page. On disturbed days the scales are multiplied by a factor, which is indicated above the panel for that day. The variations are centred on the monthly mean value, shown on the left side of the page.

1.4.2 Magnetograms

The daily magnetograms are plotted using oneminute values of D, H and Z from the fluxgate sensors, with any gaps filled using back-up data. The magnetograms are plotted to a variable scale; scale bars are shown to the right of each plot. The absolute level (the monthly mean value) is indicated on the left side of the plots.

1.4.3 Hourly Mean Value Plots

Hourly mean values of D, H and Z for the past 12 months are plotted in 27-day segments corresponding to the Bartels solar rotation number. Magnetic disturbances associated with active regions on the surface of the Sun may recur after 27 days: the same is true for geomagnetically quiet intervals. Plotting the data in this way highlights this recurrence, and also illustrates seasonal and diurnal variations throughout the year.

1.4.4 Daily and Monthly Mean Values

Daily mean values of D, H, Z and F are plotted throughout the year. In addition, a table of monthly mean values of all the geomagnetic elements is provided. These values depend on accurate specification of the fluxgate sensor baselines. Provisional and definitive values are indicated in the table as $\bf P$ or $\bf D$ respectively. It is anticipated that provisional values will not be altered by more than a few nT or tenths of arcminutes before being made definitive.

1.4.5 Geomagnetic activity indices

The Observatory K index. This summarises geomagnetic activity at an observatory by assigning a code, an integer in the range 0 to 9, to each 3-hour Universal Time (UT) interval. The index for each 3-hour UT interval is determined from the ranges in H and in D (scaled in nT), with allowance made for the regular (undisturbed) diurnal variation. The conversion from range to an index value is made using a quasi-logarithmic scale, with the scale values dependent on the geomagnetic latitude of the observatory. The K index retains the local time (LT) and seasonal dependence of activity associated with the position of the observatory.

The provisional aa index. A number of 3-hour geomagnetic indices are computed by combining K indices from networks of observatories to characterise global activity levels and to eliminate LT and seasonal effects. The simplest of these is the aa index, computed using the K indices from two approximately antipodal observatories: Hartland in the UK and Canberra in Australia. The aa index is calculated from linearisations of the Hartland and Canberra K indices, and has units of nT. The daily

mean value of aa (denoted Aa), the mean values of aa for the intervals 00-12UT and 12-24UT and the daily mean values for Hartland alone (Aa_n) and Canberra alone (Aa_s) are tabulated.

Although the *aa* index is based on data from only two observatories, provided averages over 12 hours or longer are used, the index is strongly correlated with the *ap* and *am* indices, which are derived using data from more extensive observatory networks.

The *aa* indices listed in this publication are provisional only; the definitive values are published by the International Service for Geomagnetic Indices, CRPE/CNET - CNRS, 4 Avenue de Neptune, F-94107 Saint Maur Cedex, France.

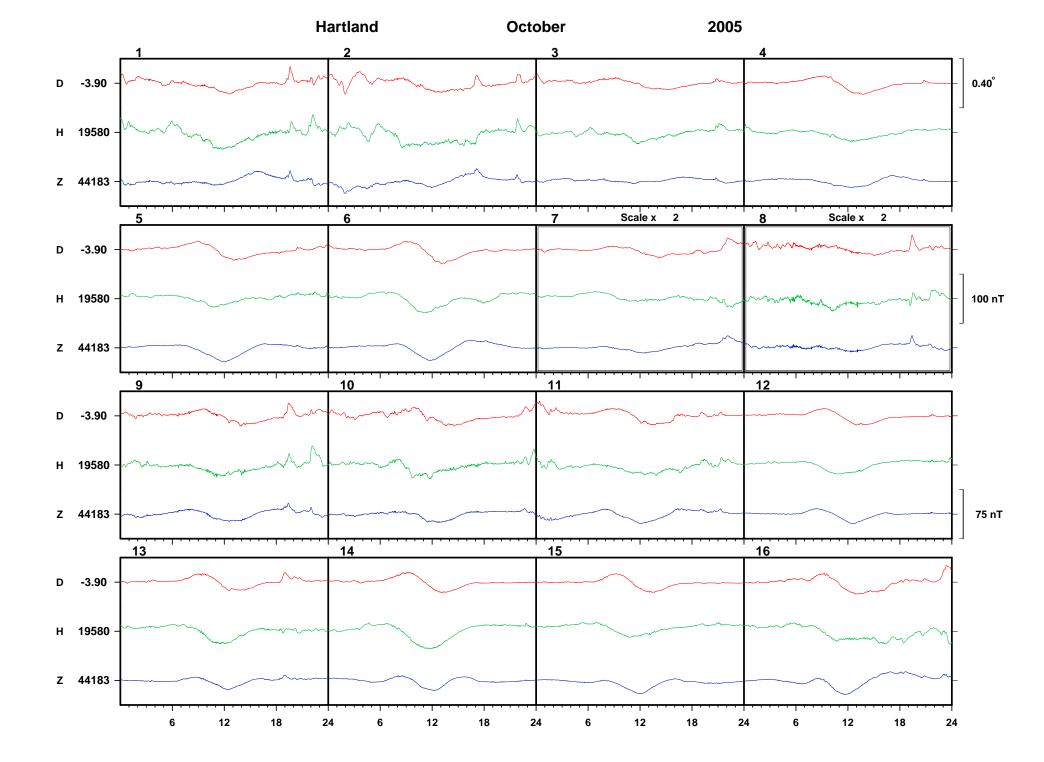
1.4.6 Rapid Variations

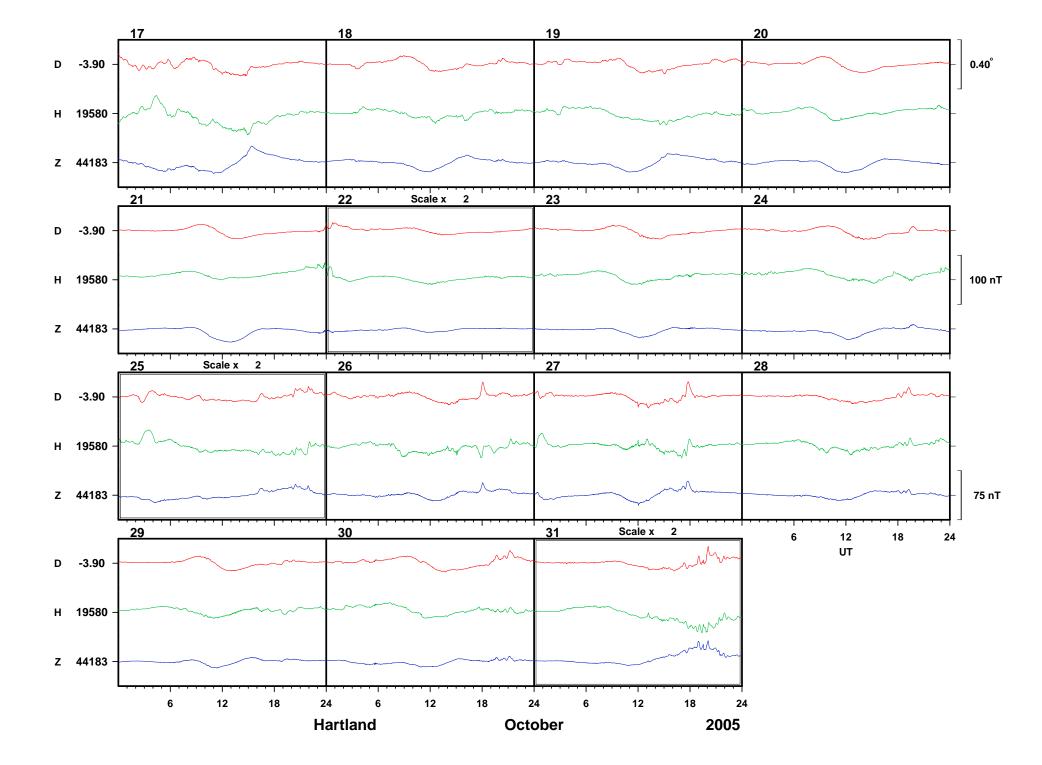
Charged particles stream from the Sun in the solar wind. The solar wind interacts with the geomagnetic field to create a cavity, the magnetosphere, in which the field is confined. When a region of enhanced velocity and/or density in the solar wind arrives at the day-side boundary of the magnetosphere (at about 10 earth radii) the boundary is pushed towards the Earth. Currents set up on the boundary of the magnetosphere can cause an abrupt change in the geomagnetic field measured on the ground and this is recorded on observatory magnetograms as a Sudden Impulse (SI). If, following an SI, there is a change in the rhythm of activity, the SI is termed a Storm Sudden Commencement (SSC). A classical magnetic storm exhibiting initial, main and recovery phases (shown by, for instance, the Dst ring current index) can often occur after a SSC, in which case the start of the storm is taken as the time of the SSC.

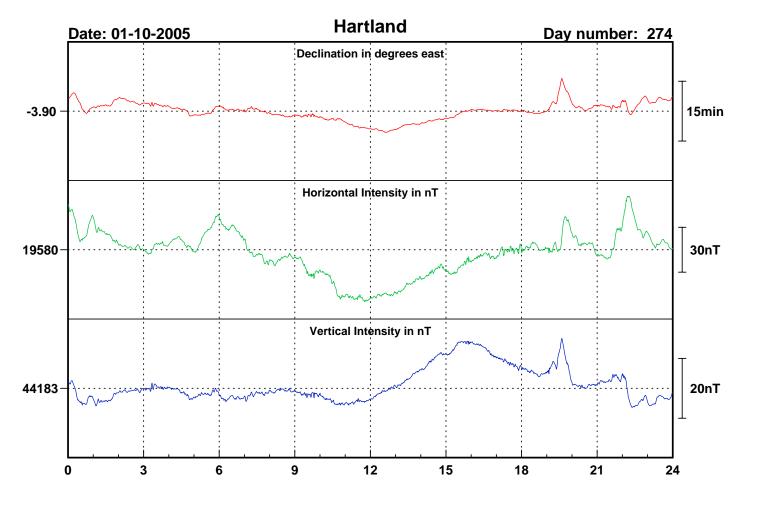
Solar flares, seen at optical wavelengths as a sudden brightening of a small region of the Sun's surface, are also responsible for increased X-ray emissions. The X-rays cause increased ionisation in the ionosphere, which leads to absorption of short-wave radio signals. On an observatory magnetogram a Solar Flare Effect (SFE), or "crochet" may be observed. This is an enhancement to the diurnal variation of the order of 10 nT, lasting about an hour.

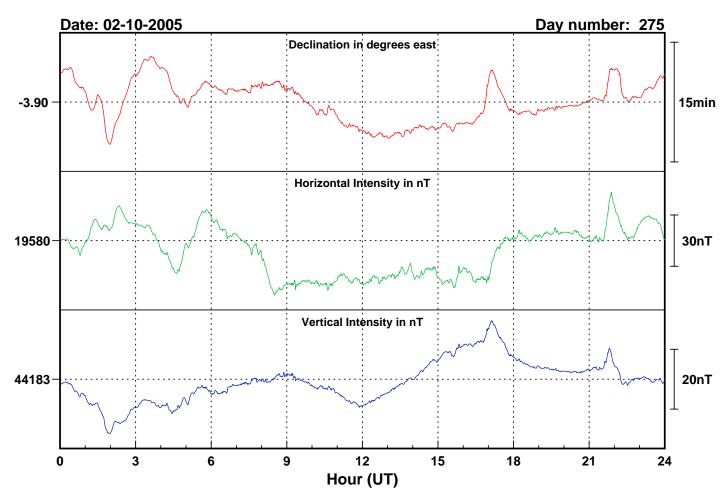
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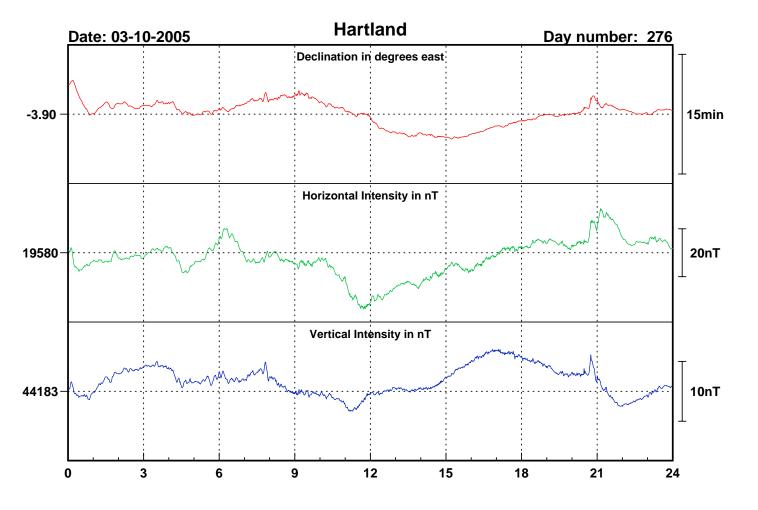
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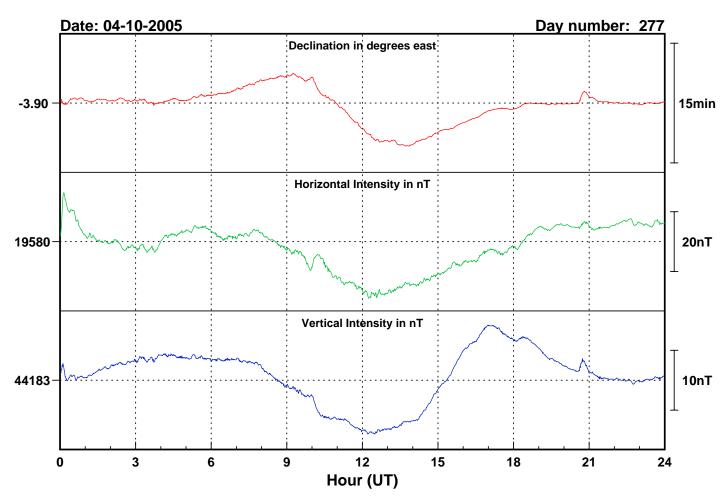


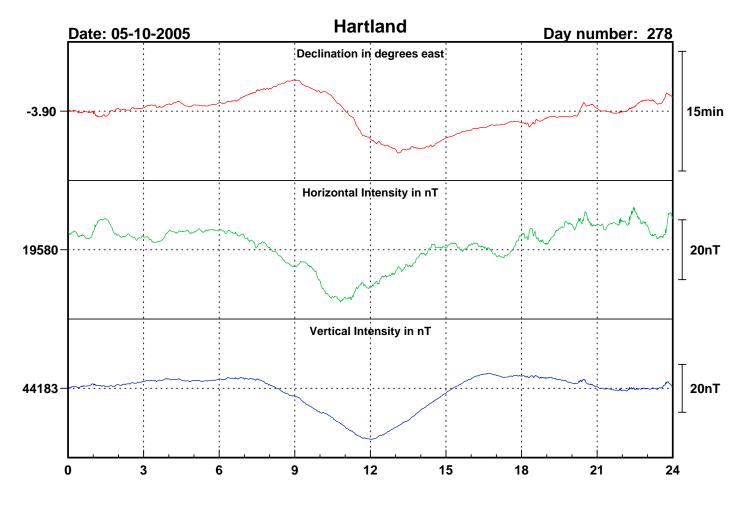


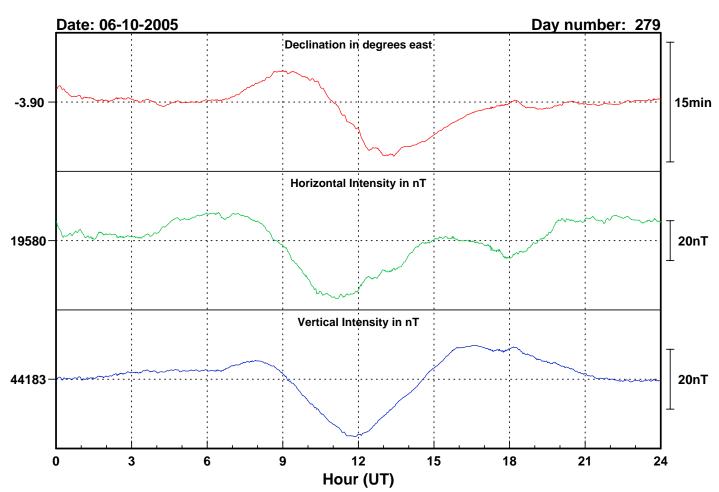


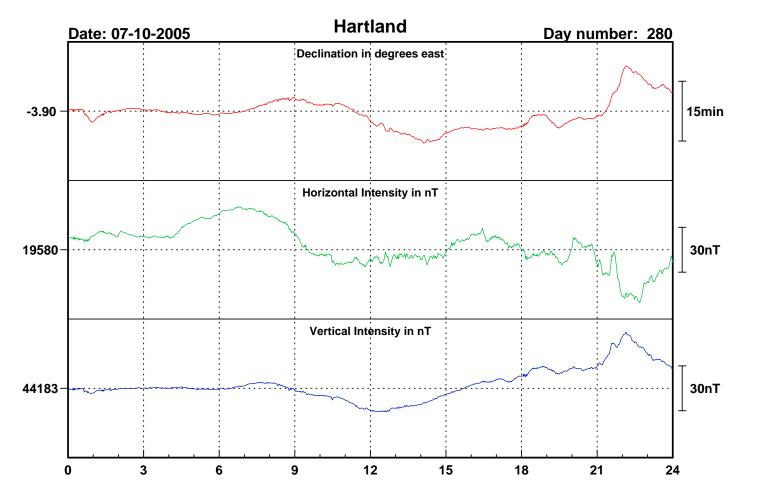


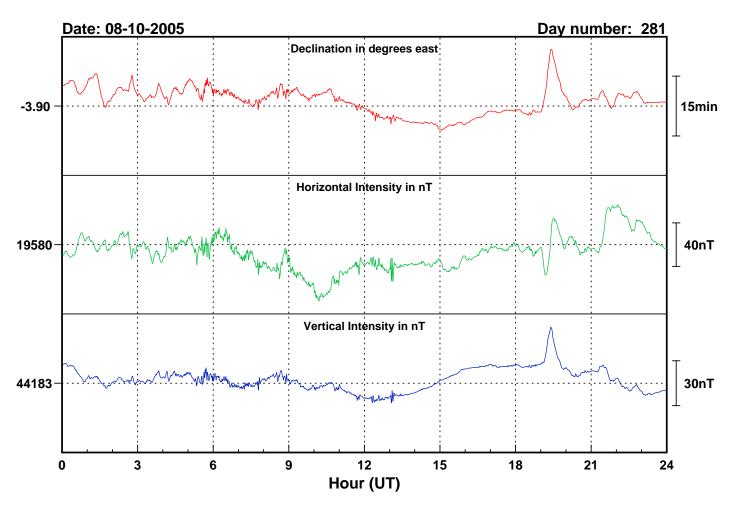


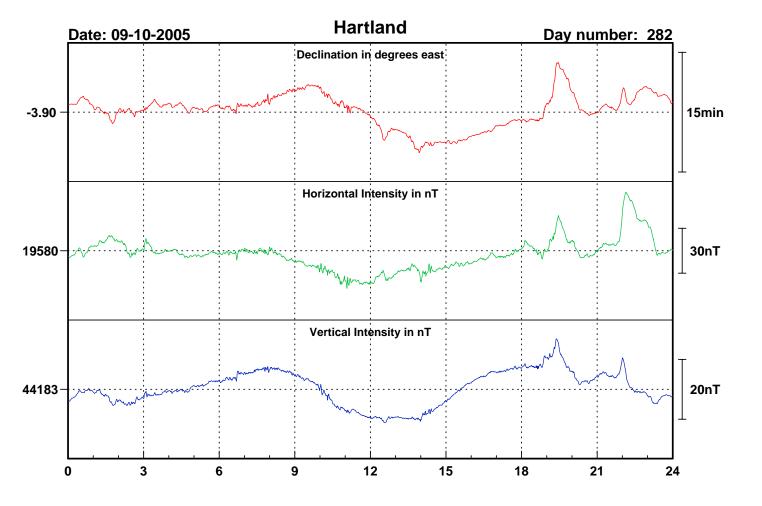


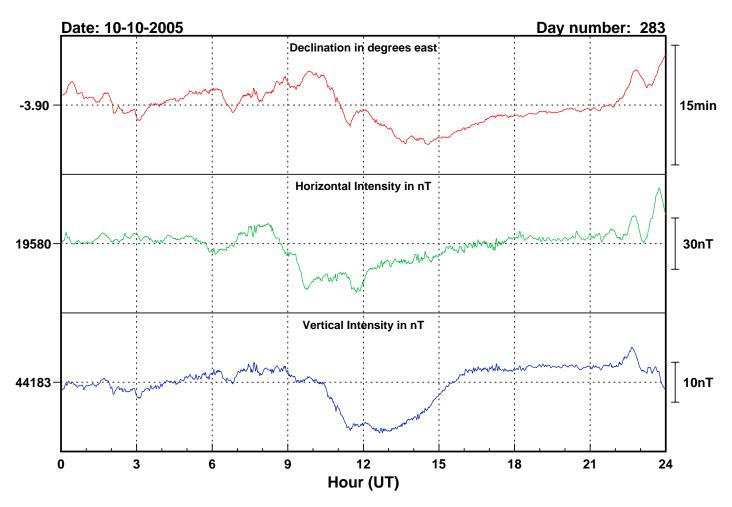


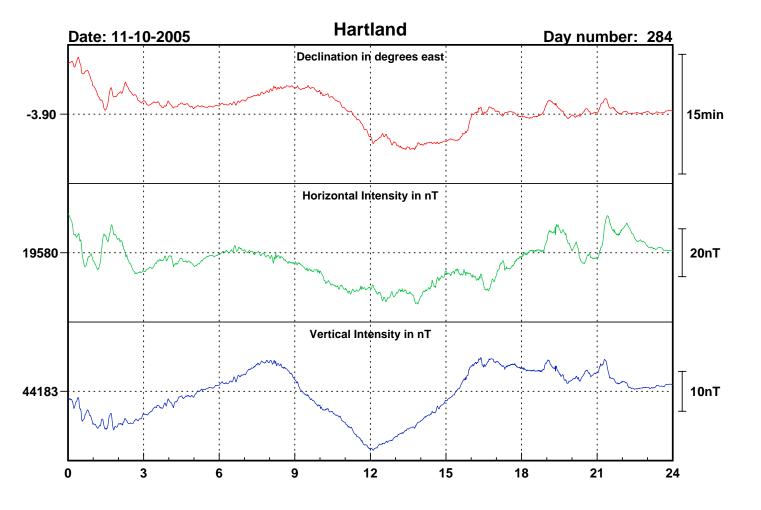


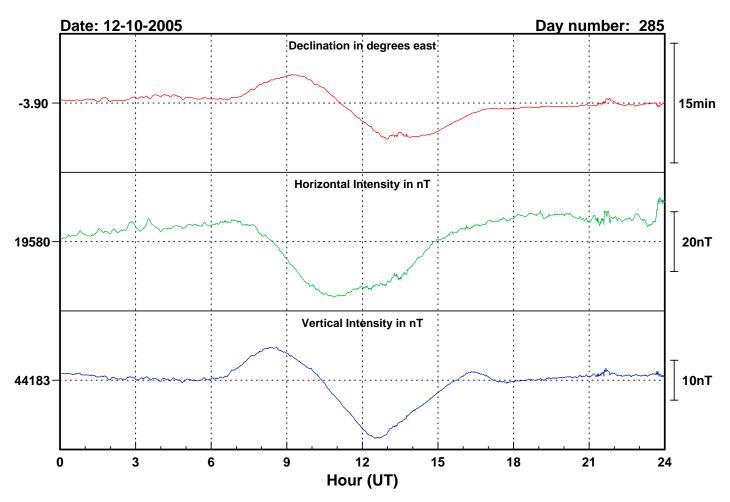


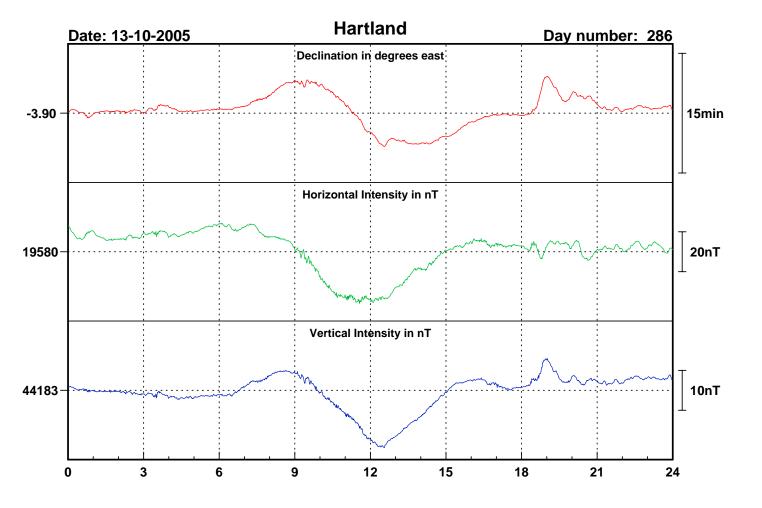


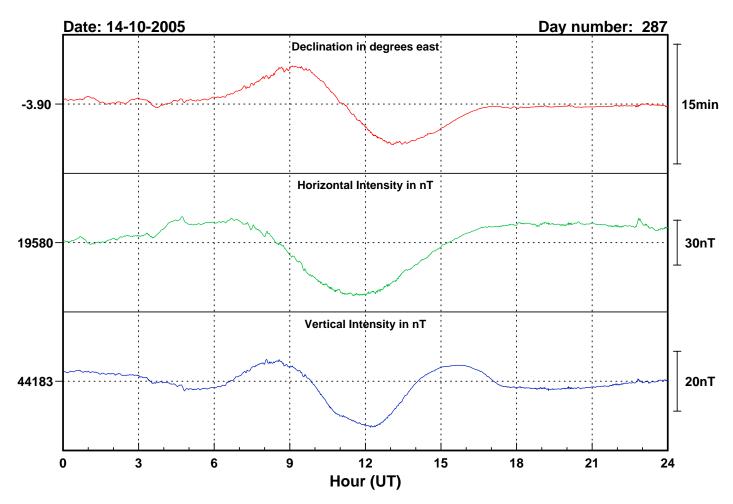


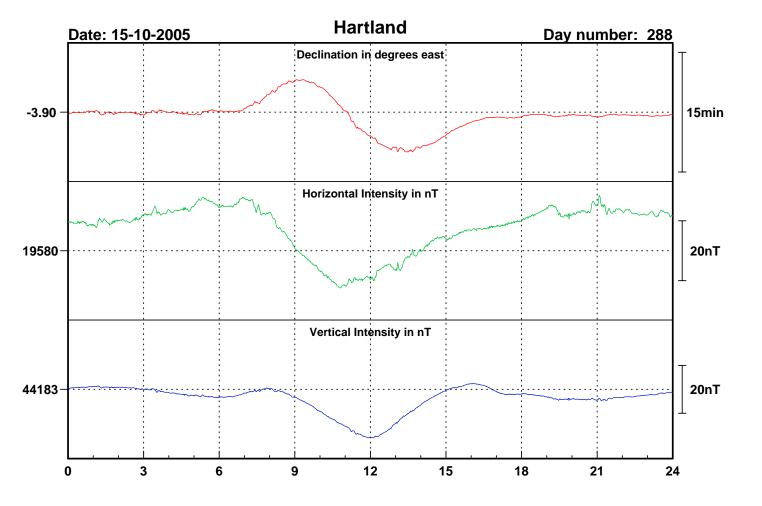


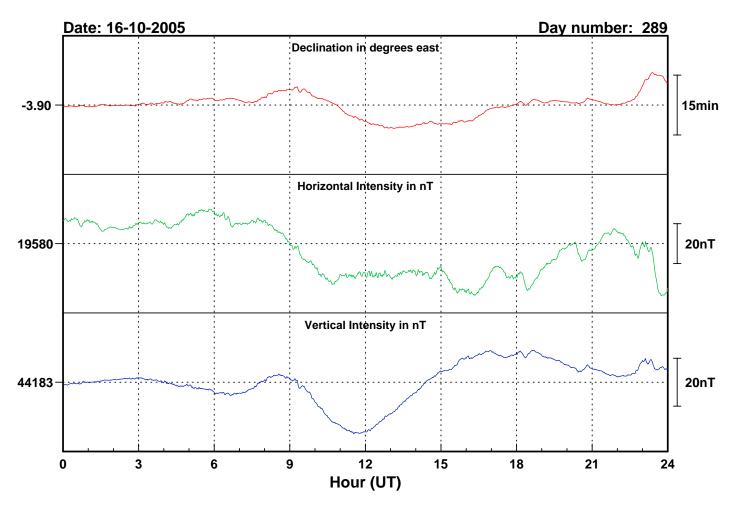


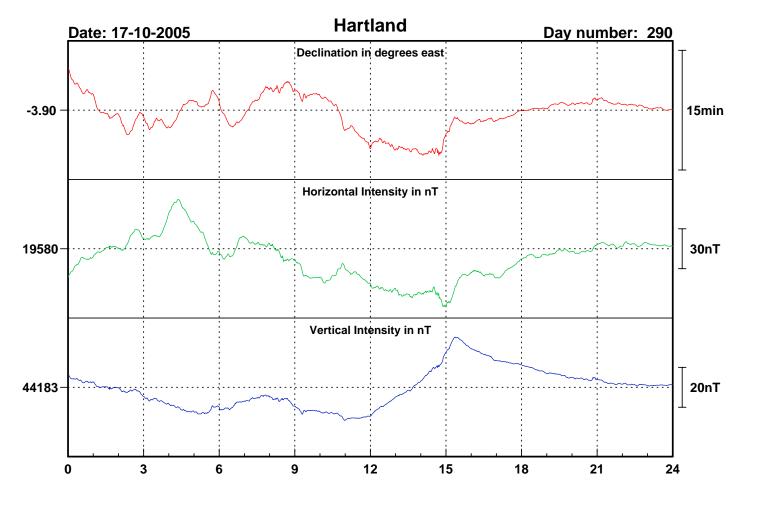


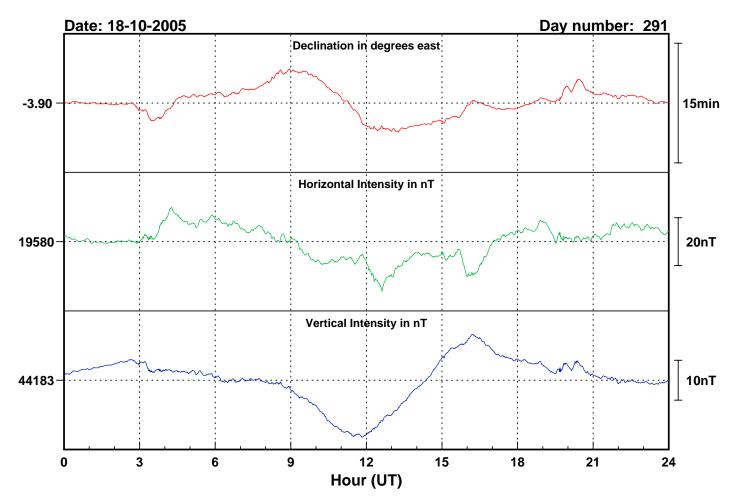


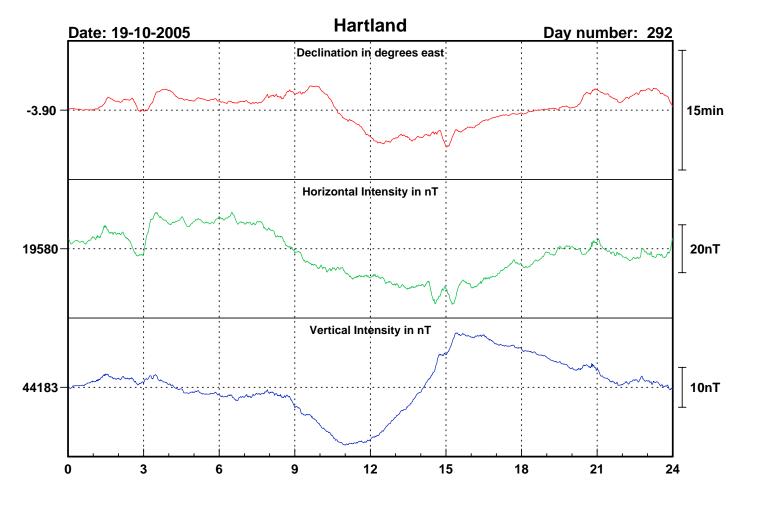


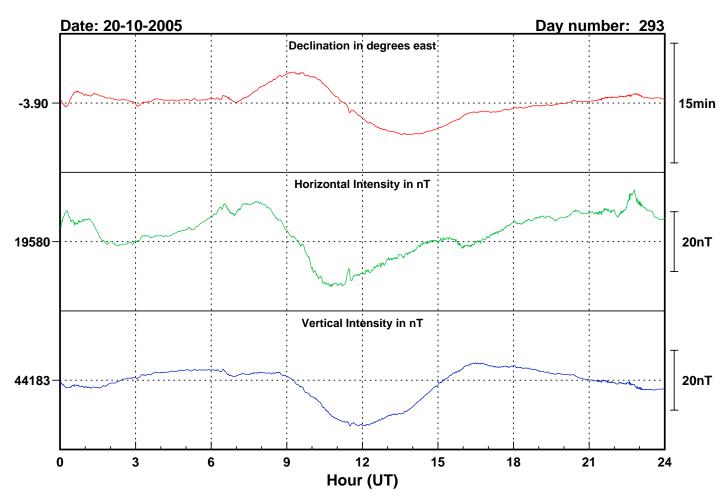


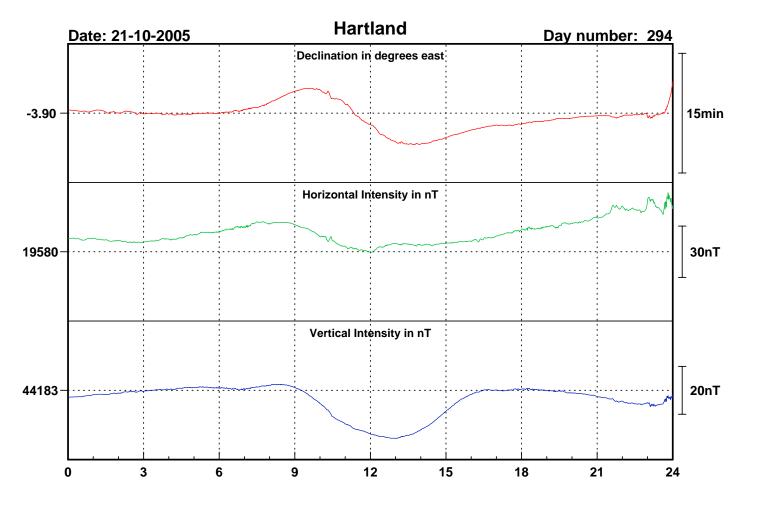


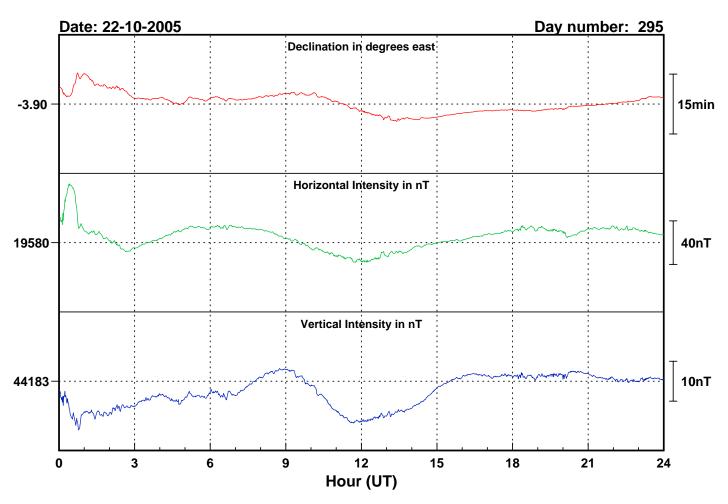


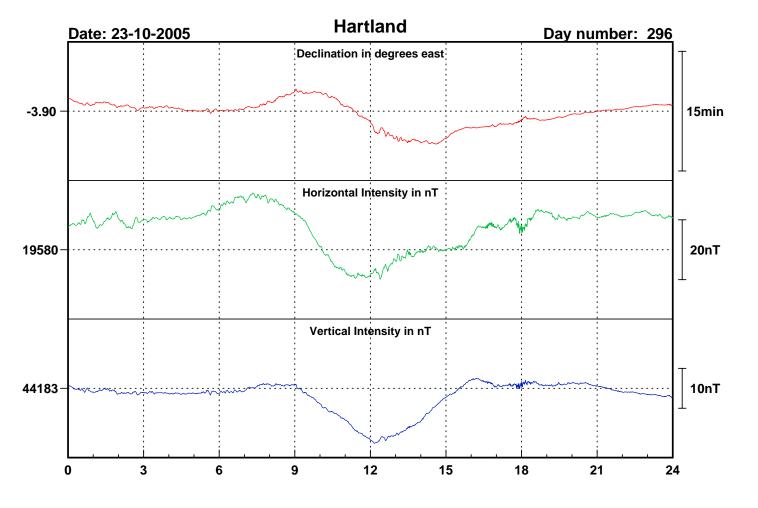


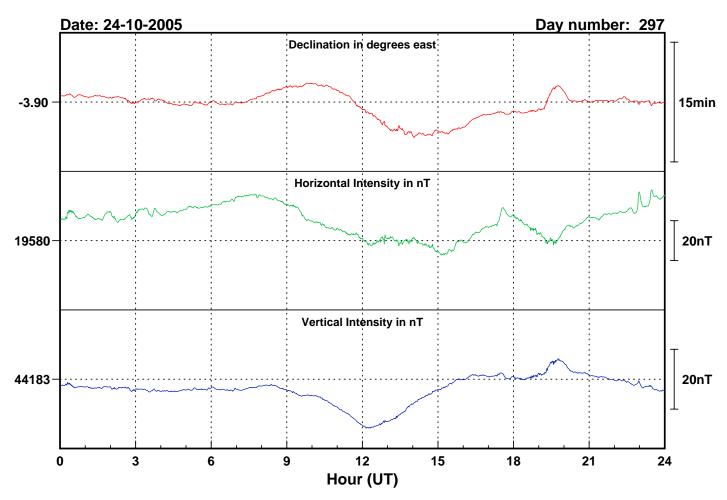


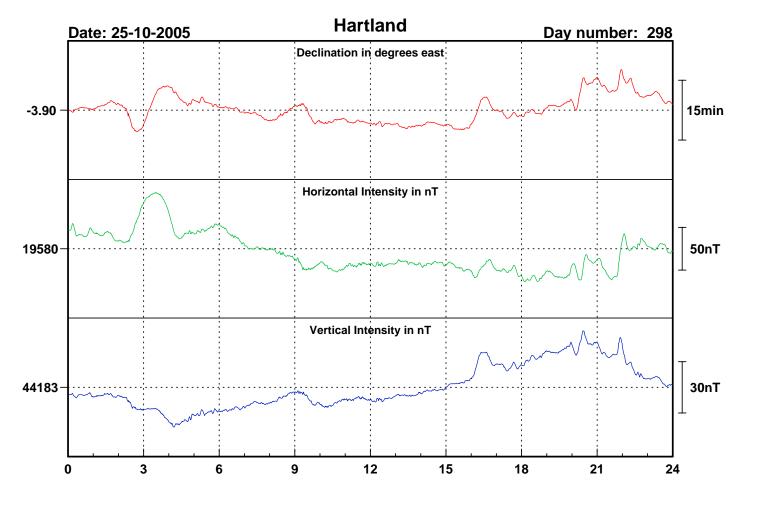


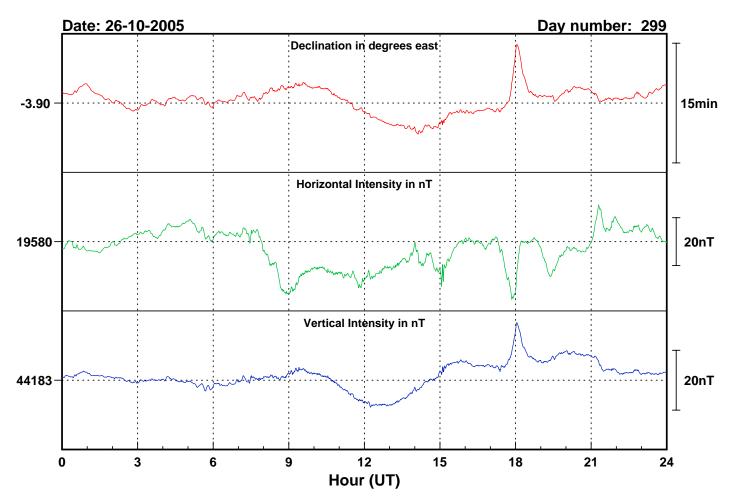


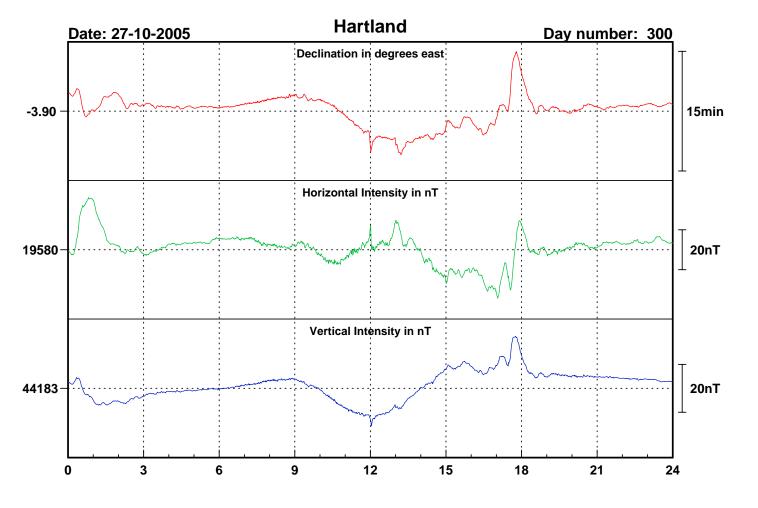


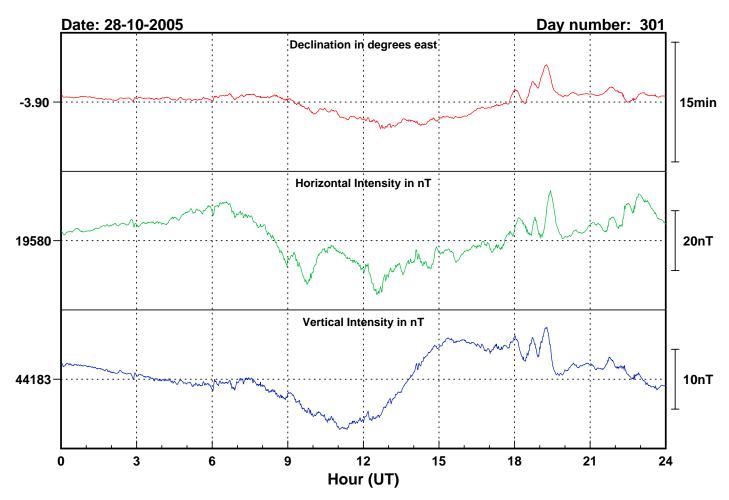


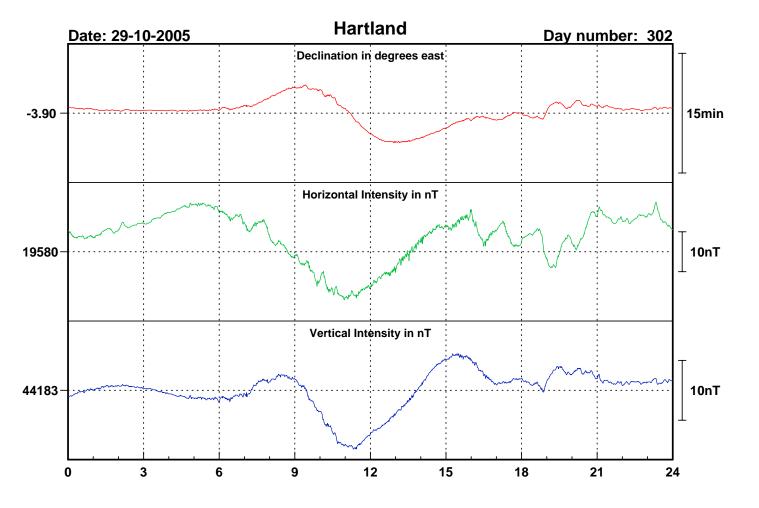


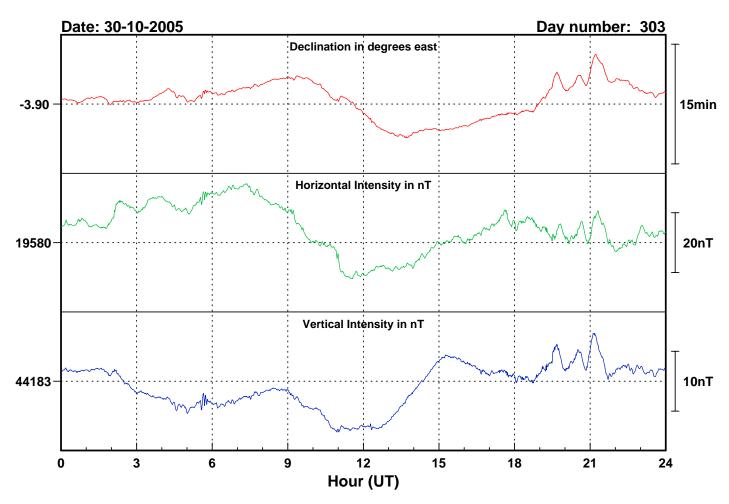


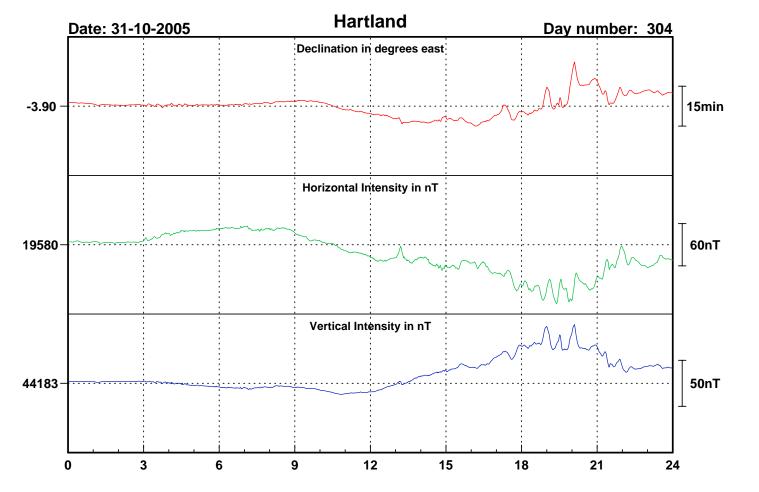


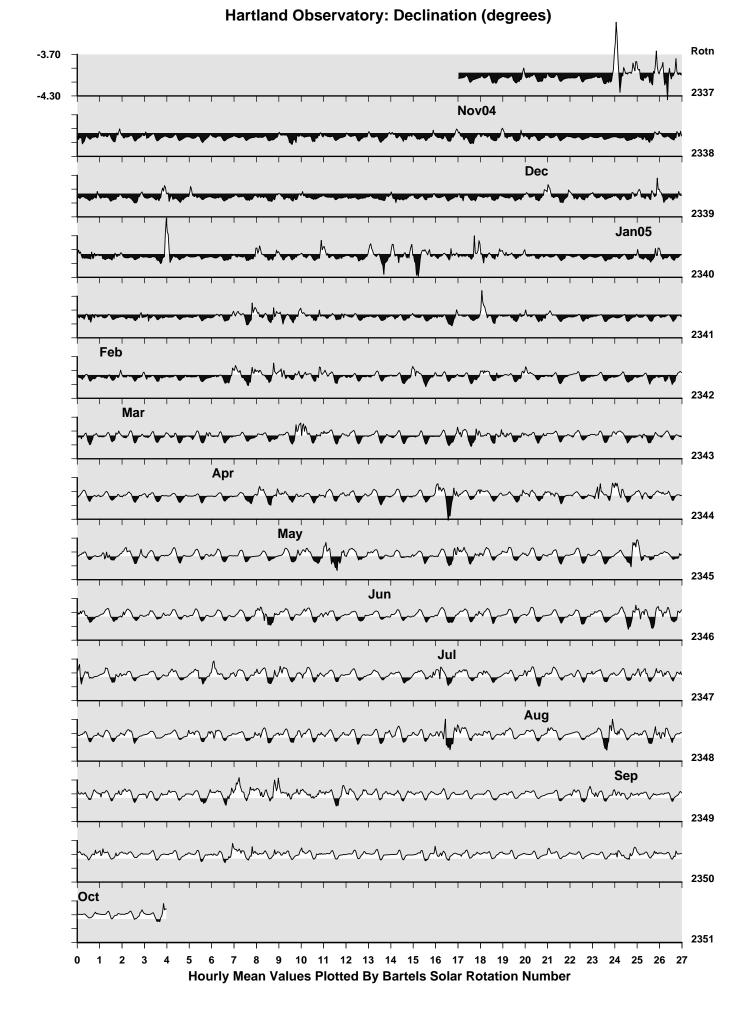




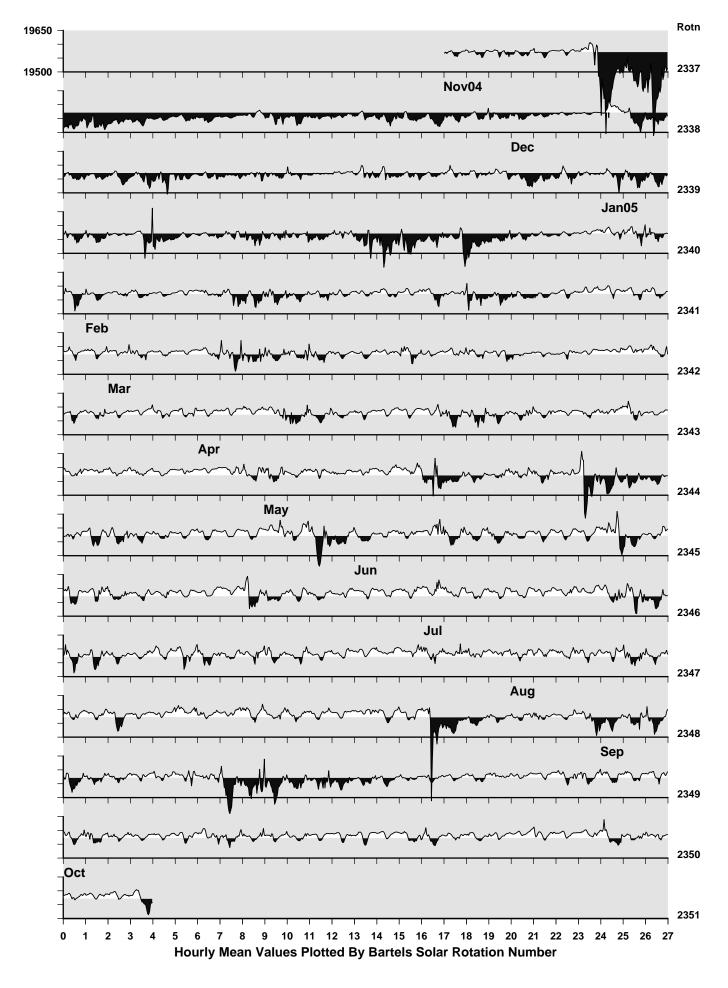




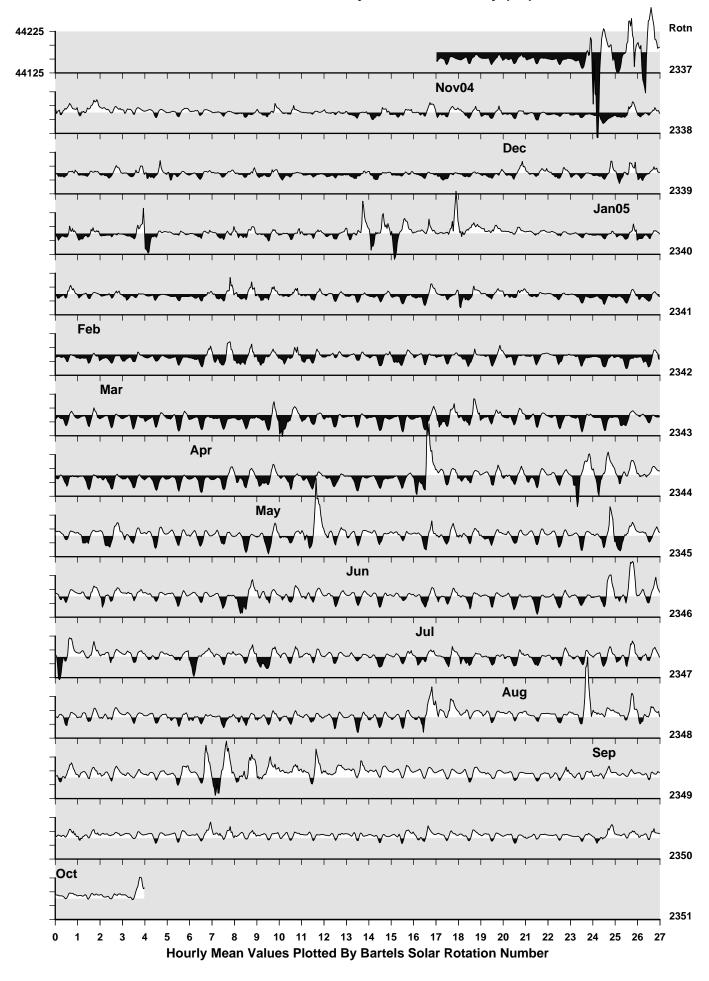


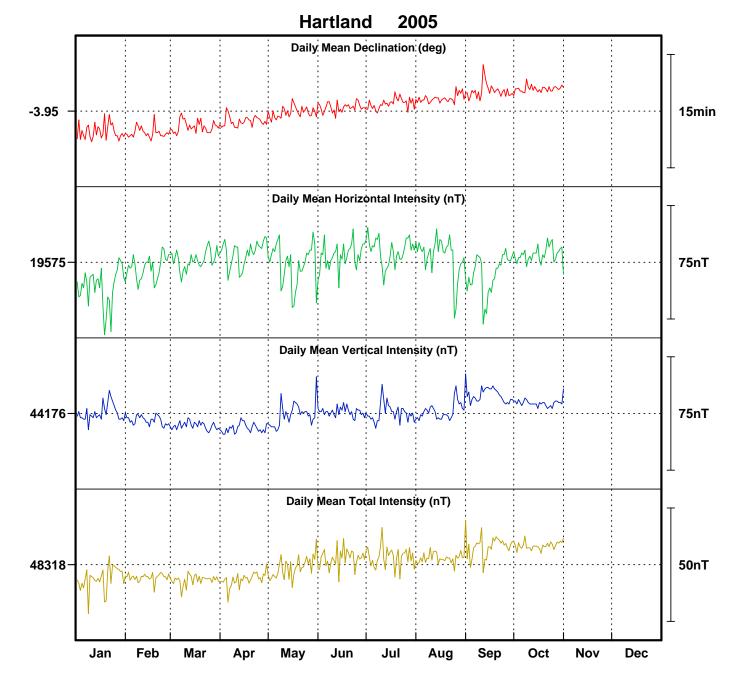


Hartland Observatory: Horizontal Intensity (nT)



Hartland Observatory: Vertical Intensity (nT)





Monthly Mean Values for Hartland Observatory 2005

Month	D	H	I	X	Y	Z	F
January	-3° 59.8′	19559 nT	66° 7.1′	19511 nT	-1363 nT	44176 nT	48312 nT
February	-3° 59.9′	19572 nT	66° 6.1′	19524 nT	-1365 nT	44171 nT	48313 nT
March	-3° 59.1′	19577 nT	66° 5.7′	19529 nT	-1360 nT	44168 nT	48312 nT
April	-3° 58.3′	19581 nT	66° 5.4′	19534 nT	-1356 nT	44166 nT	48312 nT
May	-3° 57.2′	19573 nT	66° 6.2′	19527 nT	-1349 nT	44175 nT	48317 nT
June	-3° 56.6′	19580 nT	66° 5.8′	19534 nT	-1346 nT	44177 nT	48321 nT
July	-3° 56.0′	19582 nT	66° 5.6′	19536 nT	-1343 nT	44176 nT	48322 nT
August	-3° 55.3′	19578 nT	66° 5.9′	19532 nT	-1339 nT	44178 nT	48322 nT
September	-3° 54.4′	19568 nT	66° 6.9′	19522 nT	-1333 nT	44187 nT	48326 nT
October	-3° 54.1′	19580 nT	66° 5.9′	19535 nT	-1332 nT	44183 nT	48327 nT

Note

i. The values shown here are provisional.

HARTLAND RAPID VARIATIONS

SIs and SSCs

Date	Time (UT)	Туре	Quality	H (nT)	D (min)	Z (nT)
07-11-05	10 25	SSC*	В	-4.9	-0.49	-2.0
09-11-05	06 38	SI*	C	-5.7	1.22	3.2
27-11-05	11 55	SSC*	В	6.3	-2.35	-2.8

Notes:

An asterisk (*) indicates that the principal impulse was preceded by a smaller reversed impulse. The quality of the event is classified as follows:

A = very distinct

B = fair, ordinary, but unmistakable

C = doubtful

The amplitudes given are for the first chief movement of the event.

SFEs

Date		Universal Time		H (nT)	D (min)	Z (nT)
	Start	Maximum	End			
			NONE			

Note

The amplitudes given are for the first chief movement of the event.

INDICES OF GEOMAGNETIC ACTIVITY

The K Index

Hartland Observatory October 2005

	K - INDICES FOR THREE-HOUR INTERVAL									
Day	00-03	03-06	06-09	09-12	12-15	15-18	18-21	21-24	SUM	
1	3	3	3	2	1	1	4	4	21	
2	4	3	4	2	2	4	2	3	24	
3	3	2	2	1	1	1	2	2	14	
4	2	1	0	1	1	0	1	0	6	
5	1	1	0	1	1	1	2	1	8	
6	1	1	1	1	1	1	2	1	9	
7	3	1	1	1	2	2	2	5	17	
8	4	3	4	4	3	2	5	4	29	
9	2	2	2	2	2	1	4	4	19	
10	3	2	3	3	2	1	1	3	18	
11	3	1	1	1	2	2	2	2	14	
12	0	0	0	0	1	0	0	1	2	
13	1	1	1	1	1	1	3	1	10	
14	1	1	1	0	0	0	0	1	4	
15	0	0	1	0	0	0	1	1	3	
16	1	1	2	1	1	2	2	4	14	
17	4	4	3	2	2	3	1	1	20	
18	1	2	1	1	2	2	2	1	12	
19	2	2	1	1	2	2	2	2	14	
20	2	1	1	1	0	0	0	1	6	
21	0	0	0	0	0	0	0	3	3	
22	4	2	1	1	1	0	2	1	12	
23	1	0	0	0	1	1	1	0	4	
24	1	0	0	0	1	2	2	2	8	
25	4	4	3	3	1	4	4	4	27	
26	2	1	3	2	2	4	4	3	21	
27	3	1	0	2	3	4	3	0	16	
28	0	0	1	2	2	1	3	2	11	
29	0	0	1	1	0	1	2	1	6	
30	1	2	0	1	1	1	3	3	12	
31	1	1	1	1	3	4	5	4	20	

	Lower bound (nT) for the range for each index value at Hartland Observatory										
	K-Index										
0 1 2 3 4 5 6 7 8									9		
	0 5 10 20 40 70 120 200 330 500										

The aa Index

Date	Day	K-North	K-South	(a)	(b)	(c)	(d)	(e)
01-10-05	274	3 3 3 2 1 1 4 4	22233322	30	22	24	29	26
02-10-05	275	4 3 4 2 2 4 2 3	12333311	36	21	32	25	29
03-10-05	276	3 2 2 1 1 1 2 2	12121100	15	9	15	9	12
04-10-05	277	21011010	01022110	7	8	8	7	7
05-10-05	278	11011121	00011100	8	5	5	7	6
06-10-05	279	11111121	$0\ 0\ 0\ 1\ 2\ 1\ 1\ 1$	9	7	6	10	8
07-10-05	280	3 1 1 1 2 2 2 5	0 1 1 3 2 2 3 4	25	22	13	34	24
08-10-05	281	4 3 4 4 3 2 5 4	2 4 3 4 2 2 3 3	52	33	47	38	42
09-10-05	282	2222144	21233122	25	18	17	27	22
10-10-05	283	3 2 3 3 2 1 1 3	21232102	22	14	23	13	18
11-10-05	284	3 1 1 1 2 2 2 2	21012211	15	10	11	14	13
12-10-05	285	00001001	$0\ 1\ 0\ 0\ 1\ 0\ 0\ 1$	4	5	3	5	4
13-10-05	286	11111131	11111110	11	7	8	10	9
14-10-05	287	11100001	$1\; 1\; 0\; 0\; 0\; 0\; 0\; 0$	5	4	6	3	4
15-10-05	288	00100011	$0\ 1\ 0\ 1\ 1\ 0\ 0\ 1$	4	5	4	5	5
16-10-05	289	11211224	$0\; 2\; 1\; 2\; 2\; 2\; 2\; 2$	17	13	10	20	15
17-10-05	290	4 4 3 2 2 3 1 1	13224310	28	22	30	21	25
18-10-05	291	12112221	01112211	12	9	8	13	11
19-10-05	292	22112222	21123311	14	16	12	18	15
20-10-05	293	21110001	$0\ 0\ 1\ 2\ 0\ 0\ 0\ 0$	7	5	9	3	6
21-10-05	294	00000003	00000002	6	4	2	8	5
22-10-05	295	42111021	31101011	15	10	18	8	13
23-10-05	296	10001110	11102110	5	8	5	8	6
24-10-05	297	10001222	$1\ 0\ 0\ 0\ 3\ 2\ 1\ 2$	9	11	4	16	10
25-10-05	298	44331444	3 3 3 4 2 3 1 2	45	29	42	32	37
26-10-05	299	21322443	11312221	29	14	16	28	22
27-10-05	300	31023430	11023321	23	15	12	27	19
28-10-05	301	00122132	01133111	12	13	10	16	13
29-10-05	302	00110121	10100110	7	5	5	7	6
30-10-05	303	12011133	02022112	14	11	9	16	12
31-10-05	304	11113454	0 2 2 3 4 4 4 4	35	38	12	61	37
Mo	nthly me	an value =		1	1	I		

The northern daily mean value, Aa_n The southern daily mean value, Aa_s (a)

Notes

⁽b)

The mean value of aa for the interval 00-12 UT (c)

The mean value of aa for the interval 12-24 UT (d)

The daily mean value of aa (Aa) (e)

i. The values are rounded to the nearest integer.

The units of the aa index are nT. ii.

iii. The values shown here are provisional. The definitive values are computed and published by the International Service for Geomagnetic Indices, Paris